

Stochastic Fields with Shallow Water Flow Dynamics

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Outline

- Deterministic Model
 - Physical motivation
 - Deterministic flow
- Stochastic Model
 - Covariance structures
 - 'Static' flow
- Stochastic-Deterministic Model
 - Embedding deterministic flow
 - Dynamic flow

Outline

'static' stochastic flow:
Covariance structures in
space & time

deterministic velocities:
numerical solution of
SWEs



dynamic flow:
stochastic fields with
subordinated SWE
dynamics

EXISTENCE AND SMOOTHNESS OF THE NAVIER–STOKES EQUATION

CHARLES L. FEFFERMAN

The Euler and Navier–Stokes equations describe the motion of a fluid in \mathbb{R}^n ($n = 2$ or 3). These equations are to be solved for an unknown velocity vector $u(x, t) = (u_i(x, t))_{1 \leq i \leq n} \in \mathbb{R}^n$ and pressure $p(x, t) \in \mathbb{R}$, defined for position $x \in \mathbb{R}^n$ and time $t \geq 0$. We restrict attention here to incompressible fluids filling all of \mathbb{R}^n . The Navier–Stokes equations are then given by

$$(1) \quad \frac{\partial}{\partial t} u_i + \sum_{j=1}^n u_j \frac{\partial u_i}{\partial x_j} = \nu \Delta u_i - \frac{\partial p}{\partial x_i} + f_i(x, t) \quad (x \in \mathbb{R}^n, t \geq 0),$$

$$(2) \quad \operatorname{div} u = \sum_{i=1}^n \frac{\partial u_i}{\partial x_i} = 0 \quad (x \in \mathbb{R}^n, t \geq 0)$$

with initial conditions

$$(3) \quad u(x, 0) = u^\circ(x) \quad (x \in \mathbb{R}^n).$$

Here, $u^\circ(x)$ is a given, C^∞ divergence-free vector field on \mathbb{R}^n , $f_i(x, t)$ are the components of a given, externally applied force (e.g. gravity), ν is a positive coefficient (the viscosity), and $\Delta = \sum_{i=1}^n \frac{\partial^2}{\partial x_i^2}$ is the Laplacian in the space variables. The Euler equations are equations (1), (2), (3) with ν set equal to zero.

- (A) Existence and smoothness of Navier–Stokes solutions on \mathbb{R}^3 .
- (B) Existence and smoothness of Navier–Stokes solutions in $\mathbb{R}^3/\mathbb{Z}^3$.
- (C) Breakdown of Navier–Stokes solutions on \mathbb{R}^3 .
- (D) Breakdown of Navier–Stokes Solutions on $\mathbb{R}^3/\mathbb{Z}^3$.

Fluids are important and hard to understand. There are many fascinating problems and conjectures about the behavior of solutions of the Euler and Navier–Stokes equations. (See, for instance, Bertozzi–Majda [1] or Constantin [3].) Since we don't even know whether these solutions exist, **our understanding is at a very primitive level**. Standard methods from PDE appear inadequate to settle the problem. Instead, we probably need some deep, new ideas.

Equations of motion

Conservation of momentum

- **Fluids:**

Navier-Stokes equation

$$\rho \frac{d\mathbf{v}}{dt} = -\nabla p - \rho \nabla \Phi + \nabla \cdot \mathbb{T}$$

- **Solid bodies:**

Newton's second law

$$m \frac{d\mathbf{v}}{dt} = \mathbf{F}$$

Note:

- all thermodynamic processes excluded

Shallow Water Approximation

$$w = -h\nabla \cdot \mathbf{v}_h ,$$

where

height of surface

v_h horizontal velocity

w vertical velocity

Shallow Water Equations

$$\frac{\partial u}{\partial t} = -g \frac{\partial h}{\partial x} - u \frac{\partial u}{\partial x} - v \frac{\partial u}{\partial y} + fv$$

$$\frac{\partial v}{\partial t} = -g \frac{\partial h}{\partial y} - u \frac{\partial v}{\partial x} - v \frac{\partial v}{\partial y} - fu$$

$$\frac{\partial h}{\partial t} = -\frac{\partial(hu)}{\partial x} - \frac{\partial(hv)}{\partial y}$$

- Used to model: Tsunamis, flows in rivers, internal waves, Jupiter's Atmosphere

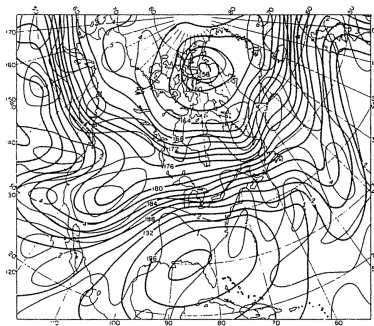
Reprinted from *Tellus*, Vol. 2, 1950, pp 237–254

Numerical Integration of the Barotropic Vorticity Equation

By J. G. CHARNEY, R. FJÖRTOFT¹, J. von NEUMANN

The Institute for Advanced Study, Princeton, New Jersey²

the Eniac. It may be of interest to remark that the computation time for a 24-hour forecast was about 24 hours, that is, we were just able to keep pace with the weather. However, much



Linear Shallow Water Equations

After linearization and reduction to one dimension.

- PDE with constant coefficients:

$$\begin{aligned}\frac{\partial u}{\partial t} &= -u_0 \frac{\partial u}{\partial x} - g \frac{\partial h}{\partial x} \\ \frac{\partial h}{\partial t} &= -u_0 \frac{\partial h}{\partial x} - h_0 \frac{\partial u}{\partial x},\end{aligned}$$

where u denotes horizontal velocity (positive towards East), h surface elevation, u_0, h_0, g constants.

One wave equation in u .

Theoretical Solution

- Solution in u :

$$u(x, t) = \pm A \sqrt{\frac{g}{h_0}} \cos(k(x - ct)) ,$$

where A denotes amplitude, k wave number and c phase velocity:

$$c = u_0 \pm \sqrt{gh_0} .$$

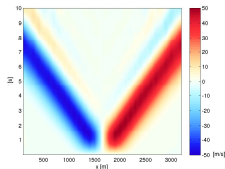


Figure: Positive (negative) velocities denote wavefronts travelling to the right (left). Notice: constant velocities along the characteristic lines $x = x_0 + (u_0 \pm \sqrt{gh_0})t$.

Numerical Solution

Consider $\alpha(i\Delta, n\Delta t) \stackrel{\text{def}}{=} \alpha_i^n$ on a tempo-spatial grid. Finite difference approximation of

- spatial derivatives

$$\left(\frac{\partial\alpha}{\partial x}\right)_i^n \approx \frac{\alpha_{i+\frac{1}{2}}^n - \alpha_{i-\frac{1}{2}}^n}{\Delta x} \stackrel{\text{def}}{=} (\delta_x\alpha)_i^n$$

- temporal derivative

$$\left(\frac{\partial\alpha}{\partial t}\right)_i^n \approx \frac{\alpha_i^{n+1} - \alpha_i^{n-1}}{2\Delta t}$$

Numerical Solution

Finite difference formulation of SWE

$$u_{i+\frac{1}{2}}^{n+1} = u_{i+\frac{1}{2}}^{n-1} - 2\Delta t \left(u_0(\overline{\delta_x u})_{i+\frac{1}{2}}^n + g(\delta_x h)_{i+\frac{1}{2}}^n \right)$$
$$h_{i+1}^{n+1} = h_{i+1}^{n-1} - 2\Delta t \left(u_0(\overline{\delta_x h})_{i+1}^n + h_0(\delta_x u)_{i+1}^n \right)$$

Consistency, accuracy and stability can be determined.

1-dim SWE simulation

particleWave.mpeg

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Stochastic Model

- 'Static' Gaussian field:

$$X(x, t) = \int_{-\infty}^t f(t-s)\Phi(x; ds)$$

- Temporal dependence

Ornstein-Uhlenbeck kernel function:

$$f(s) \propto e^{\lambda s} \mathbf{1}_{(-\infty, 0]}(s)$$

- Spatial dependence

Correlation function:

$$r_{\Phi}(x) = e^{-\frac{x^2}{2\sigma^2}}$$

- Space-time stationary covariance

$$r_X(x, t) = e^{-x^2/2\sigma^2} e^{-\lambda|t|}$$

Discrete version

- Autoregressive formulation of X^*

$$\Delta X = (\rho - 1)X + \sqrt{2(1 - \rho^2)\lambda} \Phi ,$$

where

$$\Delta X = X(x, t + \Delta t) - X(x, t)$$

$$\rho = \exp(-\lambda \Delta t)$$

$$X = X(x, t)$$

$$\Phi = \Phi_t(x) \text{ innovations with } r_\Phi(x) = e^{-\frac{x^2}{2\sigma^2}}$$

*Baxevani, A. and Podgórski, K. and Rychlik, I. (2010), Dynamically evolving Gaussian spatial fields, *Extremes*.

- $\Phi_t(x)$ is generated through its 'spectral representation'

$$\Phi(x) = \sqrt{d\omega} \sum_{j=-N}^N \sqrt{S(jd\omega)} R_j \cos(xjd\omega + \phi_j),$$

where

$S(jd\omega)$ is the discretized spectrum, corresponding to $r_\phi(x)$

$R_j \sim \text{Rayleigh}(1)$, iid, and independent of

$\phi_j \sim U[0, 2\pi]$, iid

'Static' Fields I

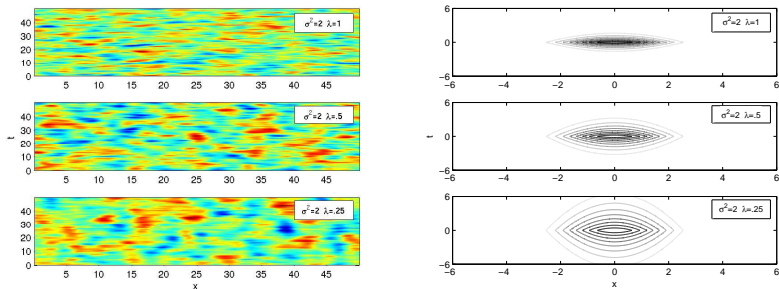


Figure: Spatio-temporal fields $X(x, t)$ for fixed σ and $\lambda = .25, .5, 1$ (left). Respective tempo-spatial covariance functions (right).

'Static' Fields II

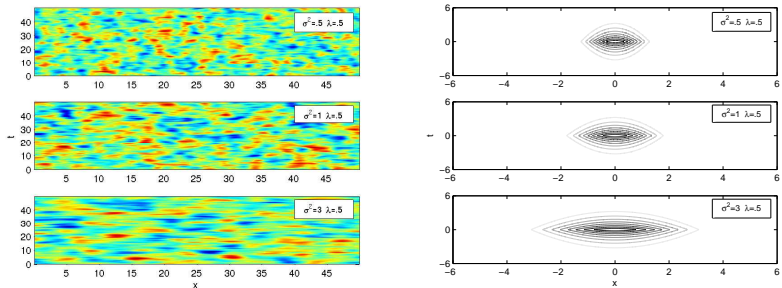


Figure: Spatio-temporal fields $X(x, t)$ for fixed λ and $\sigma^2 = .5, 1, 3$ (left). Respective tempo-spatial covariance functions (right).

Stochastic Velocities

- Velocity on random surfaces*:

$$v(x, t) = -\frac{X_t(x, t)}{X_x(x, t)} = -\frac{\frac{\partial X(x, t)}{\partial t}}{\frac{\partial X(x, t)}{\partial x}} \approx \frac{\partial x}{\partial t},$$

where

$$v \sim \text{Cauchy}(\cdot, \cdot)$$

*Baxevani, A. and Podgórski, K. and Rychlik, I. (2003).

Velocities for moving random surfaces. *Probabilistic Engineering Mechanics*.

Stochastic Velocities

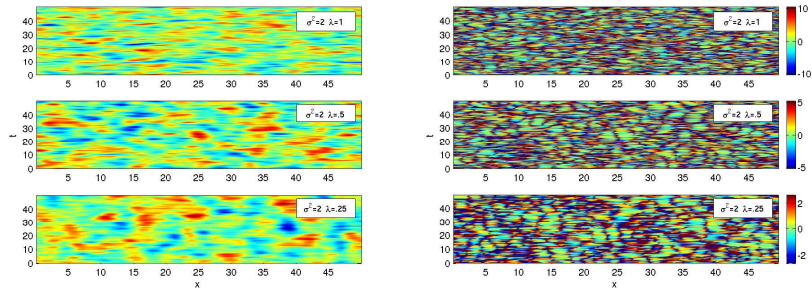


Figure: Spatio-temporal fields $X(x, t)$ (left). Respective velocities $v(x, t)$ (right). Velocities are truncated at 15% and 85% quantiles.

'No organized motion.'

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Stochastic-Deterministic Model

Embedding dynamic SWE-flow into random field $X(x, t)$

- Static field:

$$X(x, t) = \int_{-\infty}^t f(t-s)\Phi(x; ds)$$

- Deterministic flow x :

$$x_{t,h}(x) = x + \int_t^{t+h} u(x_{t,s-t}(x), s) ds$$

- Dynamic field:

$$Y(x, t) = \int_{-\infty}^t f(t-s)\Phi(x_{t,s-t}(x); ds)$$

Stochastic-Deterministic Model

- Discretized dynamic field:

$$Y(x, t + \Delta t) \approx \rho Y(x_{t, -\Delta t}(x), t) + \sqrt{2(1 - \rho^2)\lambda} \Phi_\rho(x, t)$$

'Static fields are transported according to flow x .'

Deterministic Flow

- Approximate trajectory $x_{t,s-t}(x)$:

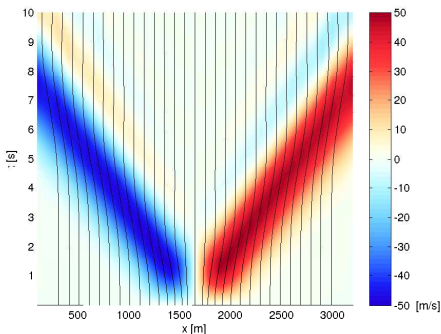


Figure: Approximate deterministic particle trajectories are given in black.

Stochastic-Dynamic Fields

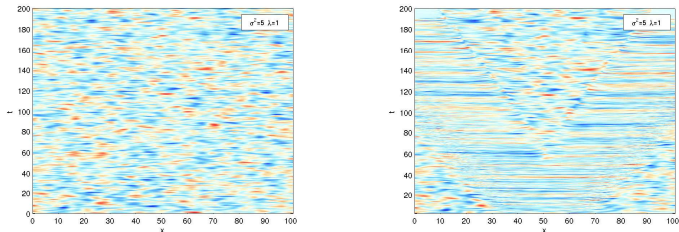


Figure: Stochastic fields for the static case (left) and with subordinated shallow water dynamics (right). Parameters are chosen to $\sigma^2 = 5$ and $\lambda = 1$.

Stochastic-Dynamic Velocities

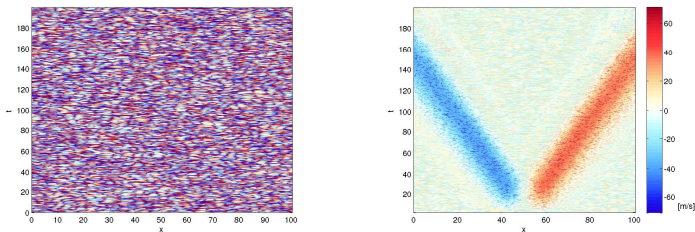


Figure: Corresponding stochastic velocities. Note that dynamic velocities accumulate along the *characteristic curves* $x = x_0 + (u_0 \pm \sqrt{gh_0})t$, whereas the static case lacks any ordered structure. Velocities are truncated.

Comparison of velocities

- Center of random velocity*, $v_c(x, t)$:

$$\frac{\int_{-\infty}^0 \left[u(x_{t,h}(x), t+h) - \frac{\partial x_{t,h}(x)}{\partial t} \right] x'_{t,h}(x) (x_{t,h}^2(x) - \sigma^2) e^{2\lambda h - \frac{x_{t,h}^2(x)}{2\sigma^2}} dh}{\int_{-\infty}^0 x'^2_{t,h}(x) (x_{t,h}^2(x) - \sigma^2) e^{2\lambda h - \frac{x_{t,h}^2(x)}{2\sigma^2}} dh},$$

where

$$x'_{t,h}(x) = \frac{\partial x_{t,h}}{\partial x}$$

*Baxevani, A. and Podgórski, K. and Rychlik, I. (2010), Dynamically evolving Gaussian spatial fields, *Extremes*.

Theorem

Assume a spatio-temporal stochastic model with the presented dependence structures. Then, the distribution of random velocity has its center at the deterministic flow velocity.

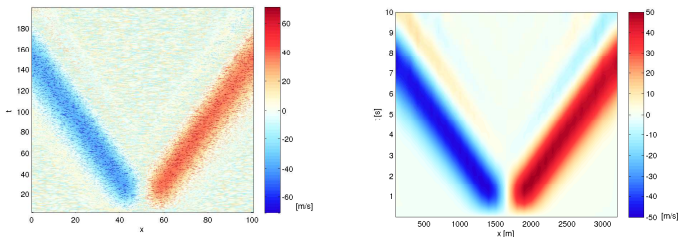


Figure: Comparison of stochastically distorted flow (*left*) to deterministic shallow water flow (*right*).

Fazit

Remarks

- different covariance structures
- higher dimensions
- non-linear dynamics

Purposes

- prediction, inference
- simple, reducing numerical burden

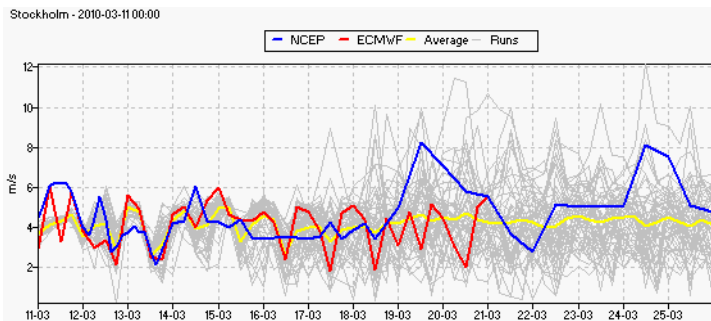


Figure: Ensemble prediction of wind (magnitude).