# Mini-course on Geometric Numerical Integration: Assignments 

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This mini-course is supported by an initiation grant from STINT
The Swedish Foundation for International Cooperation
in Research and Higher Education
February 14, 2016

## 1 Background: Ordinary differential equations and first numerical schemes

Task 1: You don't feel well after eating in a churrascaria in downtown Rio de Janeiro ... it seems that you have picked up a parasite that grows exponentially fast until treated. Returning home there are 10 of them in you. To model this problem, you should use the population dynamic model from the lecture $y^{\prime}=a y$. The growth parameter for the parasites in your body is $a=0.1+10^{-3} A$, where $A$ is the number corresponding to the last two digits of your year of birth.
a) Open an M-file called churrasc. $m$ and write a code in order to plot the exact solution of the problem for the time interval $[0,20]$. The following may be useful

```
\circ%O Code for task churrascaria
clear all
p0= ...; % initial number of parasites
tExact=[0:0.05:20]; % time interval
pExact= ...; % exact solution of the ODE
figure(),plot(tExact,pExact, ... ) % plot solution wrt time
xlabel('Time','FontSize',15) % x-axis with nice fonts
legend(...) % legend
title(...) % title
print -depsc2 -r0 taskChurrasc.eps % save the plot in .eps
```

b) Complete your M-file with an implementation of Euler's method. You should plot the numerical approximations given by the method with the step sizes $h=0.5,0.25$ and 0.1 on the interval $[0,20]$. Compare with the exact solution. The following may be useful

```
%% Code for task Churrasc
h=0.5;
t0=0; tend=20;
N=...; % compute the number of steps used by Euler's method
...; % initial time and initial value
for n=1:N
    ...; % compute one step of the method
    tEuler1(n)= ...; % store the time for the plot, see below
```

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```
```

```
pEuler1(n)= ...; % approx of pExact(t_n) for step size h
```

```
pEuler1(n)= ...; % approx of pExact(t_n) for step size h
...; % update the method
...; % update the method
end
end
h=0.25; % for the next step
h=0.25; % for the next step
...
...
% plot of the exact and numerical solutions
% plot of the exact and numerical solutions
tExact=[0:.1:20]; % time interval
tExact=[0:.1:20]; % time interval
pExact= ...; % exact solution
pExact= ...; % exact solution
figure(), plot(tExact,pExact,tEuler1,pEuler1, ... )
figure(), plot(tExact,pExact,tEuler1,pEuler1, ... )
legend(...) % legend, etc
legend(...) % legend, etc
```

...

```
```

...

```

Task 2: Consider Heun's method (1900)
\[
\begin{aligned}
Y & :=y_{0}+h f\left(x_{0}, y_{0}\right) \\
y_{1} & :=y_{0}+\frac{h}{2}\left(f\left(x_{0}, y_{0}\right)+f\left(x_{0}+h, Y\right)\right)
\end{aligned}
\]
or
\[
y_{k+1}=y_{k}+\underbrace{\frac{h}{2}\left(f\left(x_{k}, y_{k}\right)+f\left(x_{k}+h, y_{k}+h f\left(x_{k}, y_{k}\right)\right)\right)}_{=: h \phi\left(x_{k}, y_{k}, h\right)}, k=0,1,2, \ldots .
\]

Show that this numerical method has the following Butcher tableau
\begin{tabular}{c|cc}
0 & & \\
1 & 1 & \\
\hline & \(1 / 2\) & \(1 / 2\)
\end{tabular}

Show that the order of convergence of this method is \(p=2\).
Hint: Use the fact that \(f\left(x_{0}+h, y_{0}+h f\left(x_{0}, y_{0}\right)\right)=y^{\prime}\left(x_{0}\right)+h y^{\prime \prime}\left(x_{0}\right)+\mathscr{O}\left(h^{2}\right), \quad h \rightarrow 0\).

\section*{2 Geometric Numerical Integration: A taste}

Task 1: Show that
\[
I(u, v):=\ln (u)-u+2 \ln (v)-v
\]
is an invariant for the Lotka-Volterra (or predator-prey) problem
\[
\begin{align*}
& \dot{u}:=\frac{\mathrm{d} u}{\mathrm{~d} t}=u \cdot(v-2) \\
& \dot{v}:=\frac{\mathrm{d} v}{\mathrm{~d} t}=v \cdot(1-u) . \tag{1}
\end{align*}
\]

This means that \(I(u(t), v(t))=I\left(u_{0}, v_{0}\right)\) for all time \(t\) along the exact solution of our problem. Here, ( \(u_{0}, \nu_{0}\) ) are the initial values of the above problem.
Hint: Divide the two equations in (1) formally and use the method of separation of variables to find the invariant I.

Task 2: The motion of two bodies which attract each other by gravity is described by Kepler's problem
\[
\begin{align*}
& \ddot{q}_{1}=-\frac{q_{1}}{\left(q_{1}^{2}+q_{2}^{2}\right)^{3 / 2}} \\
& \ddot{q}_{2}=-\frac{q_{2}}{\left(q_{1}^{2}+q_{2}^{2}\right)^{3 / 2}} \tag{2}
\end{align*}
\]
where \(q(t):=\left(q_{1}(t), q_{2}(t)\right) \in \mathbb{R}^{2}\) denotes the position of the second body relatively to the (fixed) first body at time \(t\). This is a Hamiltonian problem with
\[
H(p, q)=\frac{1}{2}\left(p_{1}^{2}+p_{2}^{2}\right)-\frac{1}{\sqrt{q_{1}^{2}+q_{2}^{2}}}
\]
and \(p_{i}:=\dot{q}_{i}\) for \(i=1,2\). We consider the initial values
\[
q_{1}(0)=1-e, q_{2}(0)=0, p_{1}(0):=\dot{q}_{1}(0)=0, p_{2}(0):=\dot{q}_{2}(0)=\sqrt{\frac{1+e}{1-e}}
\]
where \(e:=0.6\), then the trajectory is elliptic with eccentricity \(e\) and the motion is periodic with period \(2 \pi\) (no need to prove this).
Implement the following numerical schemes (with step size \(h=2 \cdot 10^{-3}\) ) for the Kepler problem. Plot the numerical trajectories in the plane \(\left(q_{1}, q_{2}\right)\) and monitor the evolution of the Hamiltonian \(H(p, q)\) and of the angular momentum \(L(p, q):=q_{1} p_{2}-q_{2} p_{1}\) on the time interval \([0,40 \pi]\) (both quantities are invariant along the exact solution of (2)). Use the following numerical methods:
a) Euler's method
\[
\begin{aligned}
& q_{n+1}=q_{n}+h \frac{\partial H}{\partial p}\left(p_{n}, q_{n}\right) \\
& p_{n+1}=p_{n}-h \frac{\partial H}{\partial q}\left(p_{n}, q_{n}\right)
\end{aligned}
\]
b) The symplectic Euler method
\[
\begin{aligned}
& q_{n+1}=q_{n}+h \frac{\partial H}{\partial p}\left(p_{n+1}, q_{n}\right) \\
& p_{n+1}=p_{n}-h \frac{\partial H}{\partial q}\left(p_{n+1}, q_{n}\right)
\end{aligned}
\]
c) The implicit midpoint rule.

Open an M-file called taskKepler.m in order to discretise numerically Kepler's problem. The following skeleton may be useful.
```

clear all
% Initial values
e=0.6;
q1(1)=1-e; % first component of q
q2(1)=0; % second component of q
p1(1)=0;
p2(1)=sqrt((1+e)/(1-e));
h=2*10^-3; % stepsize

```

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Since the midpoint rule is implicit for the Kepler problem, one has to solve a system of implicit equations for the determination of \(q_{n+1}\) and \(p_{n+1}\). This system is of the form \(Y=G(Y)\), with \(Y:=\) \(\left(q_{n+1}, p_{n+1}\right)\), and one can use successive iterations in order to find \(Y\). Basically, starting with an initial guess for \(Y\), denoted \(Y^{(0)}\), one iterates \(Y^{(n+1)}=G\left(Y^{(n)}\right)\) until the norm of two successive approximations is below a certain tolerance, say \(10^{-6}\). The following may be useful
```

% midpoint scheme
for n = 1:N
% initial guess for p,q
p1_it=p1(n);p2_it=p2(n);
...
% first iteration
p1(n+1)=p1(n)-h*((q1 (n)+q1_it)/2)/ ...
(((q1(n)+q1_it)/2)^2+((q2(n)+q2_it)/2)^2)^(3/2);
...
% successive iterations
while norm([p1(n+1);p2(n+1);q1(n+1);q2(n+1)]- ...
[p1_it;p2_it;q1_it;q2_it])>10^-6
p1_it=p1(n+1);p2_it=p2(n+1);
% compute new values for p,q using p_it, q_it
p1(n+1)= ... ;
end
end
% Compute H and L and do the plots for implicit midpoint as above

```

\section*{3 Numerical integration of Hamiltonian systems}

Task 1: Consider Hamiltonian systems of the form
\[
\begin{equation*}
\dot{y}=f(y), \text { where } f(y)=J^{-1} \nabla H(y) \tag{3}
\end{equation*}
\]
with \(y \in \mathbb{R}^{2 d}\) and the skew-symmetric matrix \(J=\left(\begin{array}{cc}0 & I \\ -I & 0\end{array}\right)\), where \(I \in \mathbb{R}^{d \times d}\) is the identity matrix. Consider the average vector field method (AVF)
\[
y_{n+1}=y_{n}+h \int_{0}^{1} f\left(\theta y_{n+1}+(1-\theta) y_{n}\right) \mathrm{d} \theta .
\]

Using the fact that
\[
H\left(y_{n+1}\right)-H\left(y_{n}\right)=\int_{0}^{1} \frac{\mathrm{~d}}{\mathrm{~d} \theta}\left(H\left(\theta y_{n+1}+(1-\theta) y_{n}\right)\right) \mathrm{d} \theta
\]
show that the AVF method exactly conserves the energy, i. e. \(H\left(y_{n+1}\right)=H\left(y_{n}\right)\).
Task 2: Consider the Hamiltonian function \(H(y)=\frac{1}{2} p^{2}-\cos (q)\), with \(y=(p, q) \in \mathbb{R}^{2}\), in (3) and implement the AVF method using the step size \(h=2^{-6}\) and initial values \(q_{0}=1\) and \(p_{0}=0.2\). Plot the position of the pendulum \(q\) versus the time and monitor the evolution of the total energy on the time interval \([0,50]\). Since the AVF method is an implicit method, you may use the same techniques as described in Kepler's problem. Furthermore, you may use matlab command quadl in order to numerically evaluate the integrals present in the AVF method. The following may be useful
```

clear all
% initial values, time, stepsize
Qzero=1.; Pzero=0.2;
t_0=0; t_end=50;
h=2^-6; N=floor((t_end-t_0)./h);
y0=[Pzero;Qzero];
% first initial guess for y
y1t=y0;
for t=1:N
err=1;
while err > 10.^(-6)
% define the functions to integrate
yint1=@(x)( - sin(y0(2)+x.*(y1t(2)-y0(2))) );
yint2= ...
% definition of the AVF method
y1=y0+[quadl( ... )*h; ... ];
% error between two consecutive iterations
err=norm(y1-y1t,2);
% update of the iterate
y1t=y1;
end
y0=y1;y1t=y0;
% save the AVF approximations for the plots
Qavf(t)=y1(2);Pavf(t)=y1(1);
end
% plots

```

\section*{4 Highly oscillatory problems}

Task 1: Show that a trigonometric method (see Definition 4.1 in the lecture notes) satisfying
\[
\begin{aligned}
\psi(\zeta) & =\sin (\zeta) \psi_{1}(\zeta) \\
\psi_{0}(\zeta) & =\cos (\zeta) \psi_{1}(\zeta)
\end{aligned}
\]
is symplectic if and only if
\[
\psi(\zeta)=\operatorname{sinc}(\zeta) \phi(\zeta) \text { for } \zeta=h \omega .
\]

Task 2: Consider the differential equation
\[
\begin{equation*}
x^{\prime \prime}(t)=-\omega^{2} x(t), \quad \omega \gg 1 \tag{4}
\end{equation*}
\]
with initial conditions \(x(0)=x_{0}\) and \(x^{\prime}(0)=x_{0}^{\prime}\).
(a) Show, that the exact solution of the above problem is given by
\[
\binom{\omega x(t)}{x^{\prime}(t)}=\left(\begin{array}{cc}
\cos (\omega t) & \sin (\omega t) \\
-\sin (\omega t) & \cos (\omega t)
\end{array}\right)\binom{\omega x_{0}}{x_{0}^{\prime}} .
\]
(b) Consider the midpoint rule
\[
y_{n+1}=y_{n}+h f\left(\frac{y_{n}+y_{n+1}}{2}\right)
\]
for differential equations of the form \(y^{\prime}=f(y)\) as well as the Störmer-Verlet method
\[
\begin{aligned}
x_{n+1 / 2}^{\prime} & =x_{n}^{\prime}+\frac{h}{2} g\left(x_{n}\right) \\
x_{n+1} & =x_{n}+h x_{n+1 / 2}^{\prime} \\
x_{n+1}^{\prime} & =x_{n+1 / 2}^{\prime}+\frac{h}{2} g\left(x_{n+1}\right)
\end{aligned}
\]
for differential equations of the form \(x^{\prime \prime}=g(x)\). Compute one step of each of the above numerical methods applied to the problem (4). Note that for the midpoint rule, the problem needs to be rewritten as a first order system. Analyse the stability of the numerical methods with respect to the time step size \(h\). That is, write the numerical solution as
\[
\binom{\omega x_{n+1}}{x_{n+1}^{\prime}}=M(h \omega)\binom{\omega x_{n}}{x_{n}^{\prime}}
\]
and compute the eigenvalues of the matrix \(M(h \omega)\). If these eigenvalues have modulus \(\leq 1\) for all time step sizes \(h\), then the numerical method is stable.

\section*{5 GNI for SDEs}

\section*{Task 1: Simulation of a Brownian motion.}
(a) Write a Matlab code to simulate one realisation of a discretised Brownian motion \(W(t)\) on \([0,1]\). For this, one uses the fact that
\[
W\left(t_{\ell}\right)=W\left(t_{\ell-1}\right)+\mathrm{d} W_{\ell}
\]
where each \(\mathrm{d} W_{\ell}\) is an independent random variable of the form \(\sqrt{h} N(0,1)\). Consider grid points given by \(t_{\ell}=\ell h\), where \(h=2^{-4}, 2^{-6}\), and \(2^{-8}\). Plot your numerical results \(W\left(t_{\ell}\right)\) for these three different values of \(h\).
You may use the following
```

clear all
randn('state',100) % set the state of randn
% discretised Brownian motion for h=2^(-4)
Tend= ... ; h=2^(-4) ; N= ...;
W(1)= ...; % BM starts at 0 a.s.
for l=...
dW=sqrt(h)*randn(1,1); % increment/normal rand. var.
W(l+1)= ... ; % iter. procedure using def. of BM
end
% plot W against time
...
% discretised Brownian motion for h=2^(-6)
...

```
(b) With the help of the above part, compute the mean of \(W(t)\) over 200, 2000, 20000 trajectories of \(W(t)\) on \([0,1]\) with \(h=2^{-8}\). Plot your results.
```

...
dW=sqrt(h)*randn(M,N+1); % generate M samples of Wiener increments
Wmean=mean (...);
...

```
(c) Finally, in the same figure, display 5 sample paths of \(W(t)\) and the mean over 50000 trajectories of \(W(t)\) on \([0,1]\) for \(h=2^{-8}\).

\section*{Task 2: Scalar stochastic oscillators.}

Consider the system of SDEs on the time interval \([0, T]\) :
\[
\begin{aligned}
\mathrm{d} X(t) & =Y(t) \mathrm{d} t \\
\mathrm{~d} Y(t) & =-\omega^{2} X(t) \mathrm{d} t+\alpha \mathrm{d} W(t)
\end{aligned}
\]
where \(W(t)\) is a scalar Wiener process, the parameters \(\omega \gg 1\) and \(\alpha \in \mathbb{R}\), and we denote by \(X_{0}\) and \(Y_{0}\) the initial values of the problem. You may choose: \(\omega=5, \alpha=1, T=1\) for the above parameters and \(X_{0}=0, Y_{0}=1\) for the initial values.
In the lecture, we have seen that the exact solution verifies the following trace formula (i. e. linear drift in the expected energy of the harmonic oscillator)
\[
\mathbb{E}\left[\frac{1}{2} Y(t)^{2}+\frac{\omega^{2}}{2} X(t)^{2}\right]=\frac{1}{2} Y_{0}^{2}+\frac{\omega^{2}}{2} X_{0}^{2}+\frac{\alpha^{2}}{2} t \quad \text { for all time } \quad t
\]

Compute and plot the linear drift in the expected energy along the numerical solutions given by the Euler-Maruyama scheme
\[
\begin{aligned}
X_{n+1} & =X_{n}+h Y_{n} \\
Y_{n+1} & =Y_{n}-h \omega^{2} X_{n}+\alpha \Delta W_{n}
\end{aligned}
\]
where \(\Delta W_{n} \sim \sqrt{h} N(0,1)\), and the stochastic trigonometric method
\[
\begin{aligned}
X_{n+1} & =X_{n} \cos (\omega h)+Y_{n} \omega^{-1} \sin (\omega h)+\alpha \omega^{-1} \sin (\omega h) \Delta W_{n} \\
Y_{n+1} & =-X_{n} \omega \sin (\omega h)+Y_{n} \cos (\omega h)+\alpha \cos (\omega h) \Delta W_{n}
\end{aligned}
\]

The following code may be helpful:
```

clear all;randn('state',100)
% number of paths samples for
% the approximation of the expectation
M=10^3;
% Initial parameters
omega=5;X0=0;Y0=1; alpha=1;
% initial energy
Haminit= ...
Tend=5;N=2^(5); h=Tend/N;
% initialisation of matrix containing
% numerical energies
Hamem=zeros(M,N);
Hamstm=zeros(M,N);
for s=1:M % loop on the samples
% Initial values
Xtempem= ... ;Ytempem= ... ;
Xtempstm= ...;Ytempstm= ...;
for j=1:N
Winc=sqrt(h)*randn(1,N); % Wiener increment
% Euler-Maruyama
X1em=Xtempem+h*Ytempem;
Y1em= ... ;
Xtempem=X1em;Ytempem= ... ;
% energy for EM
Hamem (s,j)= ... ;
% Stoch. trigon. method
X1stm=Xtempstm* cos(h*omega)+sin (h*omega)/omega*Ytempstm+ ...
alpha*sin (h*omega)/omega*Winc;
Y1stm= ... ;
end
end
set(Q,'DefaultTextFontSize',12)
set(0,'DefaultAxesFontSize', 12)
set(0,'DefaultLineLineWidth',2)
set(0,'DefaultLineMarkerSize', 10)
% plots of the expected energy for EM, STM, exact
Dtvals=[h:h:Tend];
figure(),
plot(Dtvals,mean(Hamem),'ks-', ...
Dtvals,mean(Hamstm),'k+-', ...
Dtvals,Haminit+alpha^2.*Dtvals./2,'r'),
axis([Dtvals(1) Dtvals(end) 0 30])

```
```

hold off
xlabel('Time',' FontSize', 14)
ylabel('Energy','Rotation',0,'FontSize', 14)
legend('EM','STM','Exact')
set(gca,'FontSize',15);

```

In the above code, the expectation of a random variable \(\mathbb{E}[Z]\) is approximated by the mean \(\mathbb{E}[Z] \approx\) \(\frac{1}{M} \sum_{s=1}^{M} Z^{s}\), where \(Z^{s}\) are independent realisations of \(Z\).```

